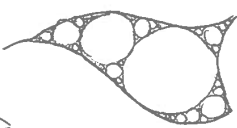


$$M = F \times$$

$$23 = 2 \cdot 3$$

$$F = \frac{M}{x}$$

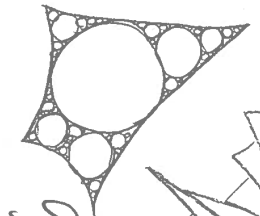
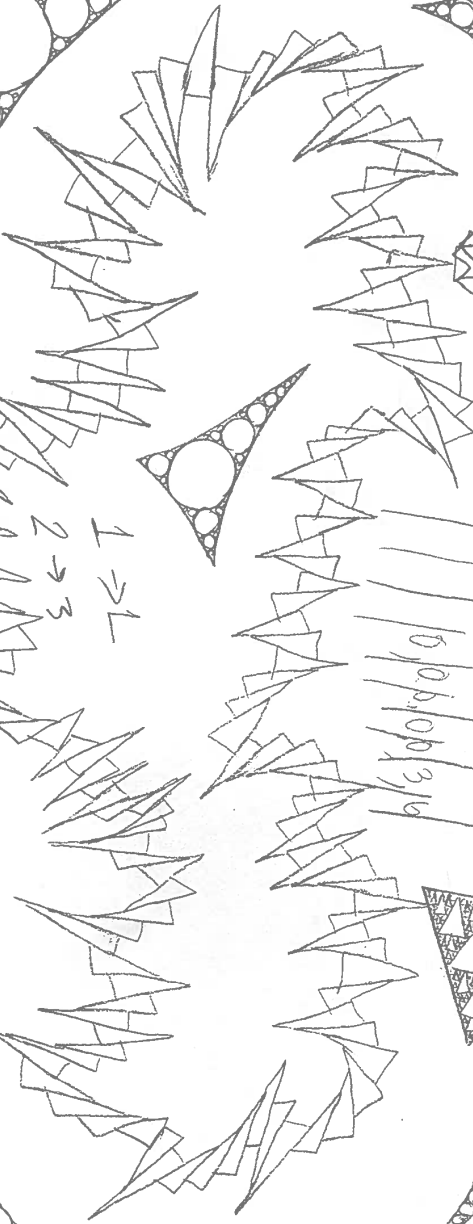
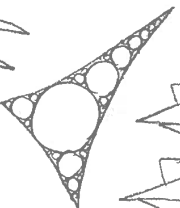
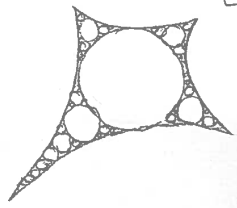
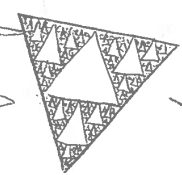
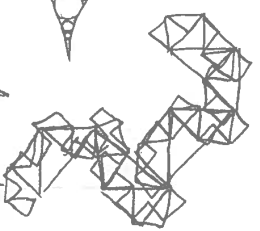
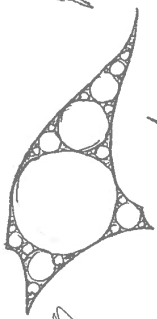
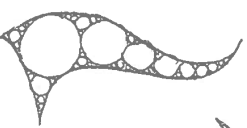
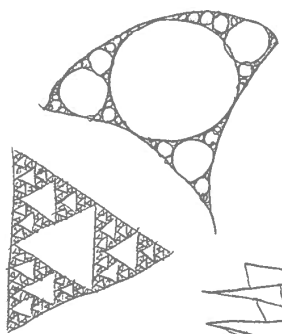
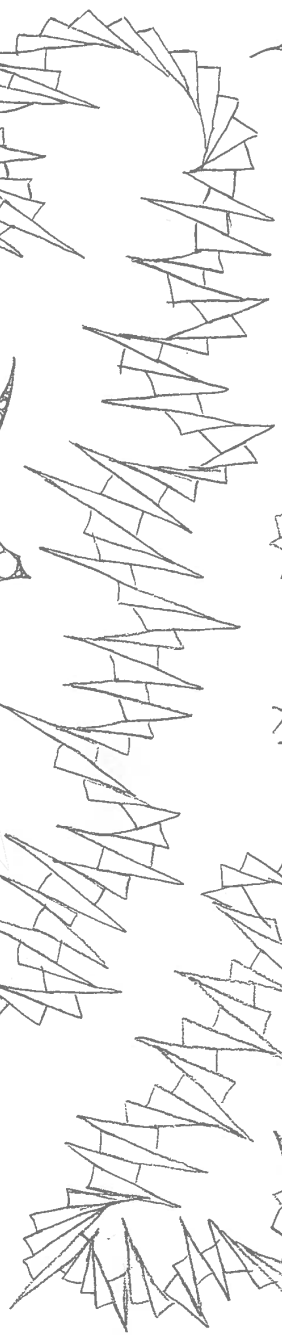


$$F = m \cdot a$$

$$a = \frac{F}{m}$$

$$a = \frac{m}{x \cdot m}$$

$$2 = \frac{100m}{5 \cdot 2 \cdot 100}$$



95-36
EJ

1 → L
2 → 3

1m diam cm
36 mm
2
10.30
10.30
13.00

24

10.30

13.00

Estonian-Finnish Olympiad 2015 Solutions

1. ANNIHILATION

i) Electron's total energy $E = \gamma m_e c^2 \implies \gamma = \frac{E}{m_e c^2}$, where $E = T + m_e c^2$ (here, $T = 1\text{MeV}$ is the given kinetic energy). Now, $\gamma = 1/\sqrt{1-v_e^2/c^2} \implies v_e = c\sqrt{1-1/\gamma^2} = c\sqrt{1-m_e^2 c^4/E^2} = \frac{c\sqrt{T^2+2Tm_e c^2}}{T+m_e c^2}$. Numerically, $v_e \approx 0.941c$.

ii) The photons fly away symmetrically with respect to the electron's trajectory: in the zero-total-momentum frame, momentum conservation implies that the photons have equal momenta and, thus, equal energies, and fly in exactly opposite directions; their energies can be equal also in the positron's frame only if they are flying totally symmetrically. Each photon gets a half of the total energy in the system: $E_\gamma = \frac{1}{2}(T + 2m_e c^2) \approx 1.01\text{MeV}$.

iii) $E_\gamma = p_\gamma c \implies p_\gamma = E_\gamma/c = 1.01\text{MeV}/c$.

iv) From $E = p_e^2 c^2 + m_e^2 c^4$, the electron's momentum is $p_e = \sqrt{\frac{E^2}{c^2} - m_e^2 c^2} = \frac{1}{c}\sqrt{T^2 + 2Tm_e c^2}$. (Equivalent result can be derived from $p_e = \gamma m_e v_e$.) Momentum conservation in the z -direction then implies that $p_e = 2p_\gamma \cos \alpha \implies \alpha = \arccos \frac{p_e}{2p_\gamma} = \arccos \frac{1}{\sqrt{1+2m_e c^2/T}} = \arctan \sqrt{\frac{2m_e c^2}{T}} \approx 45.3^\circ$.

v) In the center-of-mass frame, the total momentum of any system is zero. This means that if the outcome of a collision is only a single particle, then the particle's momentum must be zero in the center-of-mass frame. However, a photon's momentum can never be zero, because otherwise it would have zero energy and an infinite wavelength.

2. HOLOGRAPHIC LENS

i) Let $N = 0, 1, \dots$ number the zones (both opaque and transparent). The optical path

difference between two neighbouring zones must be $\lambda/2$ (opposite phase is demanded). The path difference between the N^{th} zone and the 0^{th} zone, on the other hand, is $\Delta_N = \sqrt{r_N^2 + f^2} - f$. Therefore, $\frac{N\lambda}{2} = \sqrt{r_N^2 + f^2} - f$ and $r_N = \sqrt{\left(\frac{N\lambda}{2}\right)^2 + N\lambda f}$. Only odd-numbered zones are transparent, thus we need $r_{2m+1} = \sqrt{\left(m + \frac{1}{2}\right)^2 \lambda^2 + (2m+1)\lambda f}$.

ii) A perfectly focussing glass lens is such that all the possible light rays that go to the focus have an equal optical path length. The optical path length inside a refracting medium is n times longer than the corresponding geometric length (the phase velocity is slowed down by a factor of n). Denote the sought-after thickness by x . Equate the optical path lengths of a ray through the edge of the lens and of a ray through

its centre: $\sqrt{\left(\frac{d}{2}\right)^2 + f^2} = f - x + nx \implies x = \frac{1}{n-1} \left[\sqrt{\left(\frac{d}{2}\right)^2 + f^2} - f \right] \approx 2.4\text{cm}$.

iii) Firstly, note that the given pulse is short enough that the whole lens never illuminates the focus — the pulse is only $\frac{rc}{\lambda} = 18$ periods long, but $r_{2 \times 18} \approx 1.3\text{mm} \ll 5\text{cm}$. This implies that only a thin strip of the lens is illuminating the focus at a time. The intensity, when the N^{th} period is being observed, is proportional to the area of the N^{th} zone. This is $A_N = \pi(r_{N+1}^2 - r_N^2) = \pi\left(\frac{N\lambda^2}{2} + \frac{\lambda^2}{4} + \lambda f\right)$.

As N is proportional to time (the period of the wave is constant), the intensity will also grow linearly in time. The linear part starts at $N = 1$ with a jump and ends at $N_{\text{max}} \frac{\lambda}{2} = \sqrt{\left(\frac{d}{2}\right)^2 + f^2} - f$ with a jump back into darkness, when the light from the edge of the lens arrives. The total duration of illumination is (approximately) $\tau_{\text{hol}} = N_{\text{max}} \frac{\lambda}{c} = \frac{2}{c} \left[\sqrt{\left(\frac{d}{2}\right)^2 + f^2} - f \right] \approx 7.9 \times 10^{-9}\text{s}$.

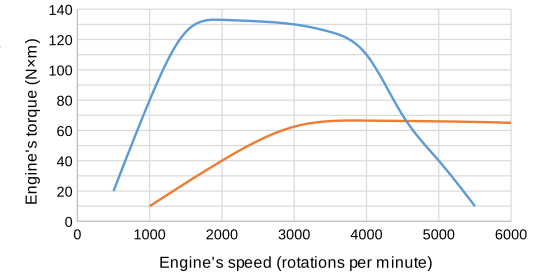
iv) The pulse is localized into a region of space with a width $\Delta x = c\tau$. Because of the Heisenberg's uncertainty principle, the pulse is composed of photons with a range of momenta (if we take the picture that the properties of the individual photons are classical) or, from a different viewpoint, is a single photon with a somewhat uncertain momentum; either way, the characteristic width in the momentum space is $\Delta p = \frac{\hbar}{\Delta x} = \frac{\hbar}{c\tau}$. The wavelength of a photon, whose momentum p is known, is $\lambda = \frac{\hbar}{p}$ (this is the de Broglie relation: the photon's energy is pc and also $h\nu = \frac{hc}{\lambda}$). Thus, $\Delta \lambda \approx \left| \frac{d}{dp} \frac{\hbar}{p} \right| \Delta p = \frac{\hbar \Delta p}{p^2} = \frac{h \frac{\hbar}{c\tau}}{\left(\frac{\hbar}{\lambda}\right)^2} = \frac{\lambda^2}{2\pi c\tau} \approx 4.4 \times 10^{-9}\text{m}$.

v) The spread of the arrival times of waves with different wavelengths is the largest for the waves that spend the longest time inside the lens. Therefore it is enough to consider only the waves that go through the thickest part of the lens — its centre. The spread in the arrival times is $\Delta t = \Delta \frac{x}{v_g} = \frac{x \Delta v_g}{v_g^2} = \frac{x \Delta \lambda}{v_g^2} \frac{dv_g}{d\lambda}$. To find the group velocity v_g itself, we can use the hint (given during the examination) that in this question we may assume the group velocity and the phase velocity to be equal (in reality it would be an unusual coincidence): $v_g = v_p$ and the phase velocity $v_p = c/n$. Therefore, $\Delta t = \frac{x \Delta \lambda}{v_g^2} \times 0.02 \frac{v_g}{\lambda} = 0.02 \frac{x n \Delta \lambda}{c \lambda} \approx 2.1 \times 10^{-14}\text{s}$. The total observed pulse length behind the glass lens is $\tau_{\text{gl}} = \tau + \Delta t = 5.1 \times 10^{-14}\text{s}$.

3. GEAR SHIFT

i) The acceleration of the car is proportional to the torque applied to the wheels. Therefore we should keep the torque onto the wheels as big as possible, and shift gears when the torque applied onto the wheels in the second gear is bigger than the one obtained in the first gear. At the same car's speed, the engine's speed is $(14 : 1) : (7 : 1) = 2$ times bigger

in the first gear than in the second gear. Correspondingly, at the same engine's speed, the torque onto the wheels is two times smaller in the second gear than in the first gear. Therefore we can draw another graph where the engine's torque is twice as small and the engine's speed is two times bigger. The intersection of those graphs is the point where the gear should be changed.



From the graph we read that at that moment in the first gear, the engine's torque is $\tau_e = 66.5\text{N}\cdot\text{m}$ and the engine's speed is $\omega_e = 4550\text{rpm}$. The wheels' angular speed is then $\omega_w = \frac{1}{14}\omega_e$ and the car's speed is $v = \frac{1}{2}\omega_w d = \frac{1}{28}\omega_e d$. Let's convert $1\text{rpm} = 120\pi\text{rad/h}$ and $d = 6 \times 10^{-4}\text{km}$. Thus, $v \approx 37\text{km/h}$.

ii) At the optimum point, the acceleration before and after gear change is the same. From Newton's second law, it equals $a = F/m = \frac{2\tau_w}{md}$, where the net torque applied to the wheels is $\tau_w = 14\tau_e$. Hence, $a = \frac{28\tau_e}{md} \approx 2.2\text{m/s}^2$.

4. STAR WARS

i) The period of a Keplerian orbit having major semi-axis a can be expressed using relation $T^2 = \frac{4\pi^2 a^3}{GM}$. However, if we didn't know the formula, we should recall Kepler's 3rd law $T^2 \sim a^3$ and simply express the period for a circular orbit with $a = r$. Therefore, $T^2 = \frac{4\pi^2 r^2}{v^2}$, where we need to insert v from $\frac{v^2}{r} = \frac{GM}{r^2}$.

For finding the period, we can get the major semi-axis from the expression for orbit's

energy decreases and our constant current source dissipates some energy (because our electromotive forces work against it). In total, the applied mechanical power $Fv = \frac{dE_m}{dx}v + I(\mathcal{E}_1 + \mathcal{E}_2)$, from where $F = -\frac{B^2 A_2}{\mu_0} + \frac{2IBA_2 N}{l} = \frac{\mu_0 I^2 N^2 A_2}{l^2}$.

10. VAPOUR PRESSURE

We construct a manometer by fixing the pipe into a “U” shape using the stand, and filling the tube partially with water. With a small trouble we should get the water to

the bottom of “U” shape and then we can measure the pressure difference from differences of depth $\Delta p = \rho g \Delta h$. After we attach a bottle to one end, we should let the pressures equalise through a needle hole in the bottle. We then squirt the unknown liquid to the bottle and close the needle hole with tape as fast as possible. We then shake the bottle to hasten the vaporisation and write down the difference of depths Δh after it has reached a stable value. During all this we should be careful to heat the bottle with our body as little as possible.

If the volume would have been fixed we would get the vapour pressure directly from the manometer reading as according to Dalton’s law. But since the diameter of the pipe was not that small we should take the relative volume increase to account.

If the depth difference is Δh , we have a relative increase of pressure $n_p = \frac{\rho g \Delta h}{p_0}$ and a relative increase of volume $n_V = \frac{\Delta h}{2V_0}$. Assuming an ideal gas, this gives us relative molar increase $n_n = \frac{(1+n_p)(1+n_V)p_0 V_0 - p_0 V_0}{p_0 V_0} \approx n_p + n_V$. Substituting in the given values we

get $n_n \approx 0.0987 \text{ m}^{-1} \times h + 0.00471 \text{ m}^{-1} \times h \approx 0.1034 \text{ m}^{-1} \times h$.

Since all the molar increase is due to the vapour, we can get the vapour pressure (partial pressure exerted by vapour) by multiplying the pressure with the molar fraction of vapour $p_x = (p_0 + \Delta p) \frac{n_n}{1+n_n}$.

The unknown liquid was ethanol with vapour pressure of $p_x = 6.52 \text{ kPa}$ at 21.6°C . However, the grading scheme was not insistent on the exact value as it proved technically quite difficult to get it correct.